

**International Summer
School of Science
Bakuriani, 2012**



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Energy Sources of the Luminous Universe

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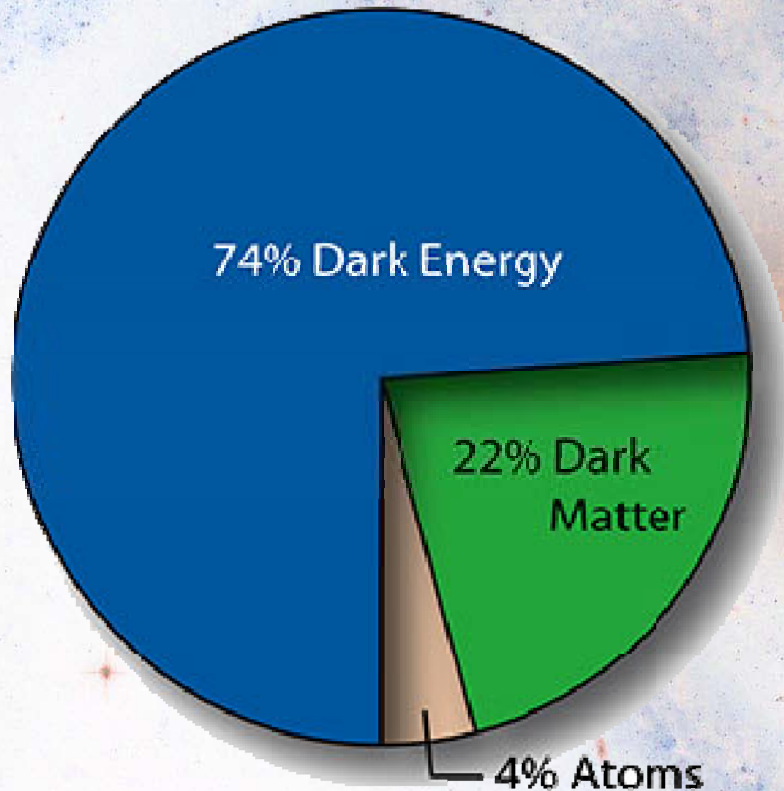
Tbilisi State University, Tbilisi, Georgia



05 July 2012

Composition of the Universe

Wilkinson MAP data (cosmological observation)



4% Atoms =

3.5% Interstellar Matter

(Hydrogen, Helium, neutrino)

0.5% Luminous matter

(Stars, Galaxies)

Studying the Luminous Universe:

Reconstruct the total picture using only 0.5%

Visible Universe



Clouds? go online: www.google.com/sky/

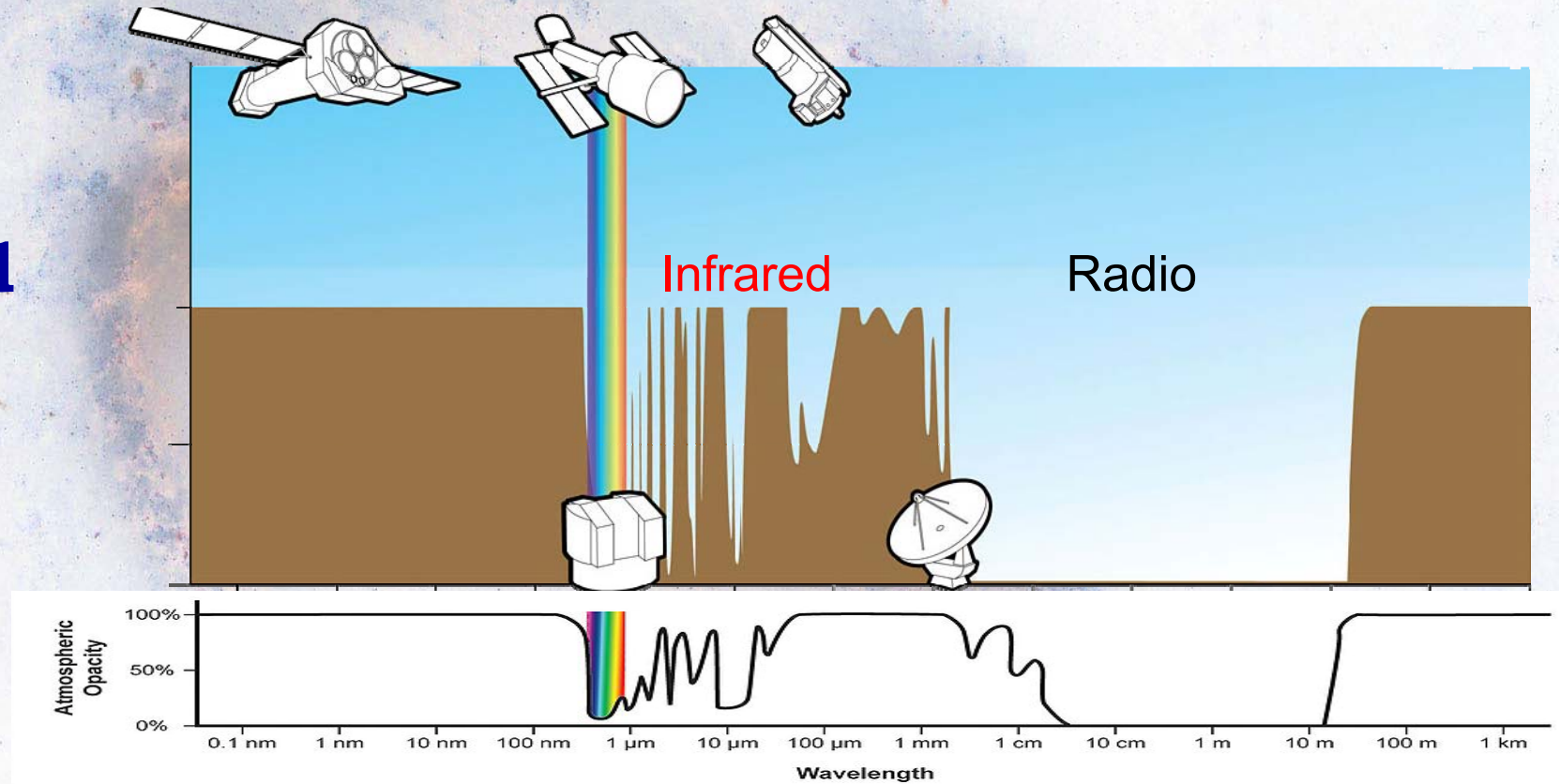
Observing Universe with naked eye

International Summer School of Science, Bakuriani 2012

Atmospheric Opacity

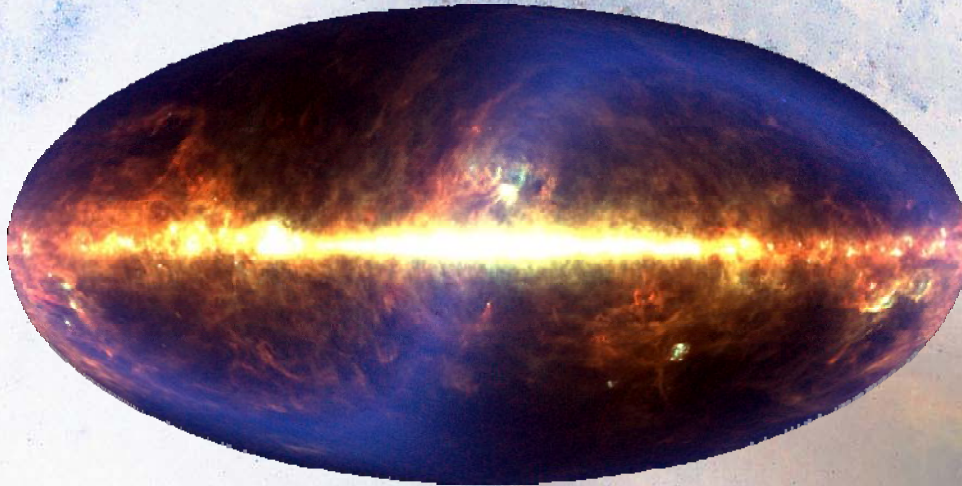
Earth Atmosphere has only few transparent windows to observe the Universe above us:

- **Optical**
- **Infrared**
- **Radio**

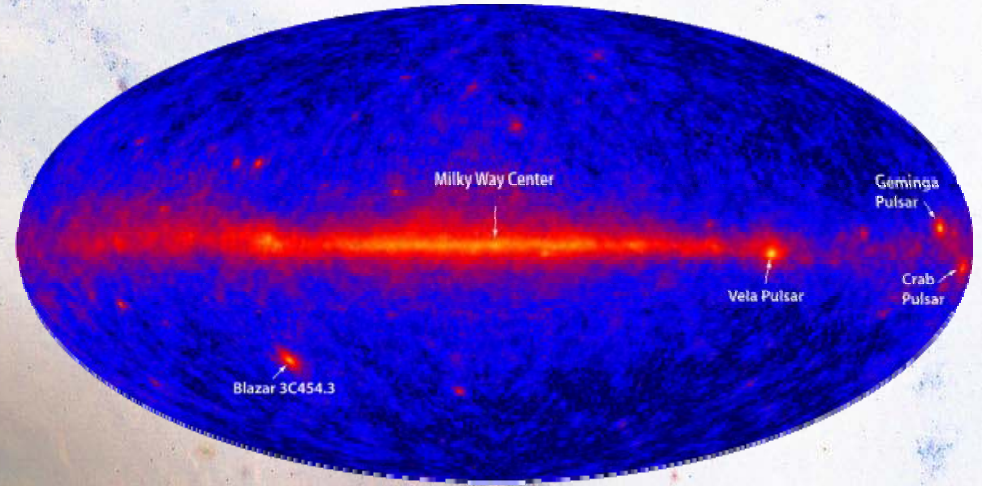


Satellite observations: re-discovery of the Universe

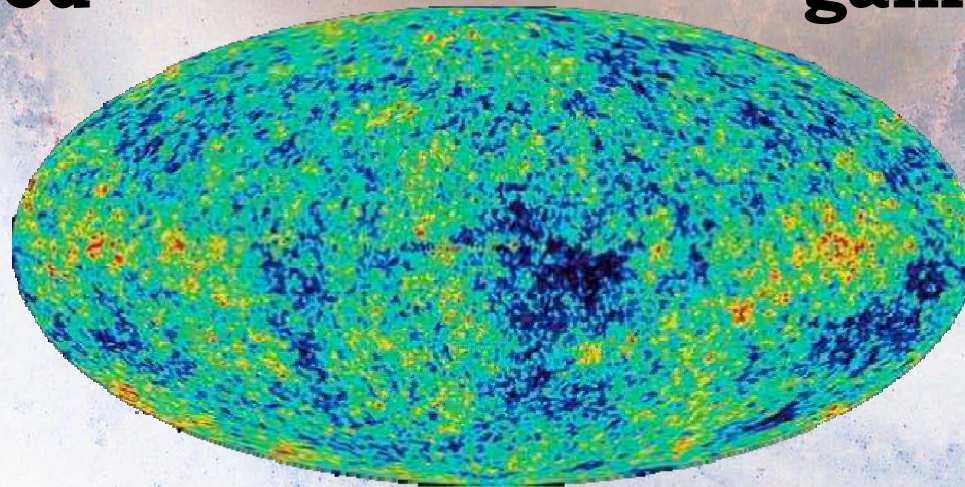
Non-Optical Universe



Infrared



gamma-ray



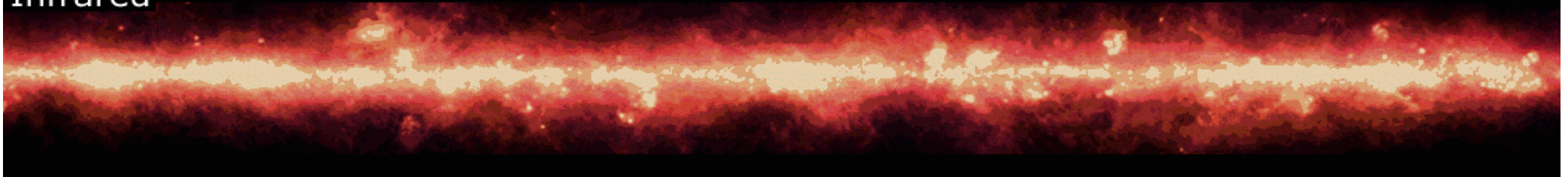
Microwave

Milky Way

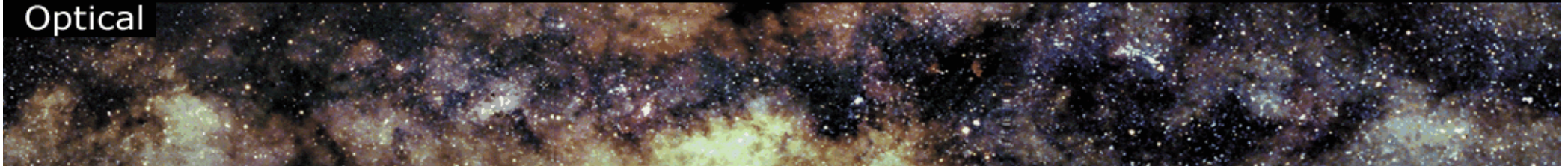
Milky way images in
different wavelength:
Infared, Optical, Gamma-ray



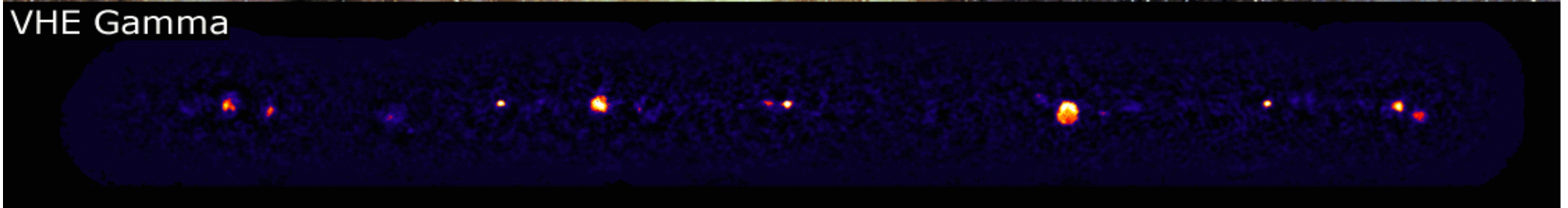
Infrared



Optical



VHE Gamma



Mass-Energy Equivalence

**Mass of the body is the measure
of its energy content**



Energy Content

Energy in wooden stick: **1kg**

$$E_0 = mc^2 = 1 (300\,000\,000)^2 = 9 \cdot 10^{16} \text{ J}$$

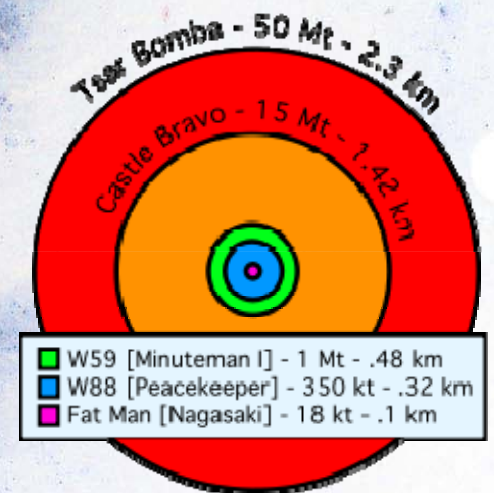
c – speed of light

TNT equivalent: 1 Ton = $4 \cdot 10^9$ J

$$E_0(1\text{kg}) = 22.5 \text{ megaton TNT}$$

Energy content of the 1kg of wood is the half of the energy released during biggest H-bomb explosion on the Earth
(x1000 Nagasaki explosions)

How to release this energy?



Energy Release

Energy production methods:

Method 1. Chemical Reactions

(e.g. burning wood)

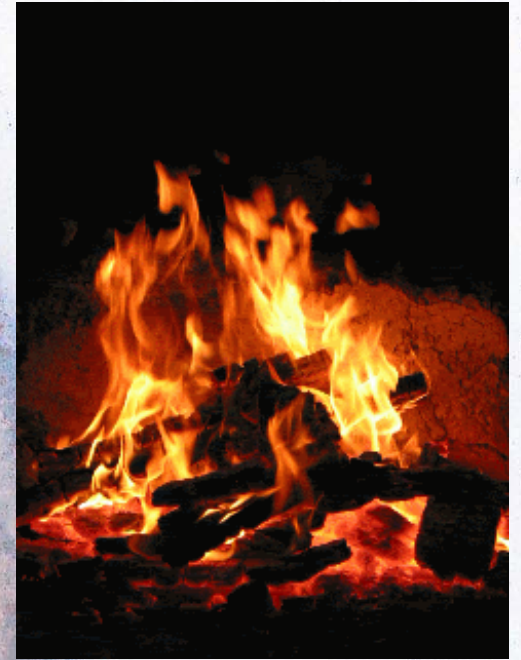
$$E_{\text{chem}} = \lambda m = 20 \text{ MJ/kg } 1 \text{ kg} = 2 \cdot 10^7 \text{ J}$$

λ – Heat of combustion

$$(E_0 = 9 \cdot 10^{16} \text{ J})$$

$$E_{\text{chem}} / E_0 = 2.22 \cdot 10^{-10} \approx 0$$

**Energy release of the wood after oxygen-burning
is negligible (0.00000002% mc^2)**



Nuclear Energy

Masses before and after reaction do not match:

$$m_1 \neq m_2 \quad (\Delta m = m_2 - m_1)$$

Energy Production: **$E = \Delta m c^2$**

Method 2. Nuclear Fission

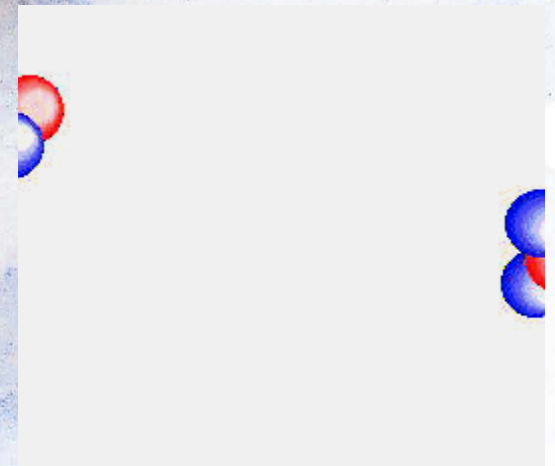
$$E \sim 0.1\% \quad (0.001 mc^2)$$

- Atomic bomb (U^{235})
- Nuclear Reactors ☺

Method 3. Nuclear Fusion

$$E \sim 0.7\% \quad (0.007 mc^2)$$

- H-Bomb (H^1, H^2, H^3)
- Controlled Nuclear Fusion ☹

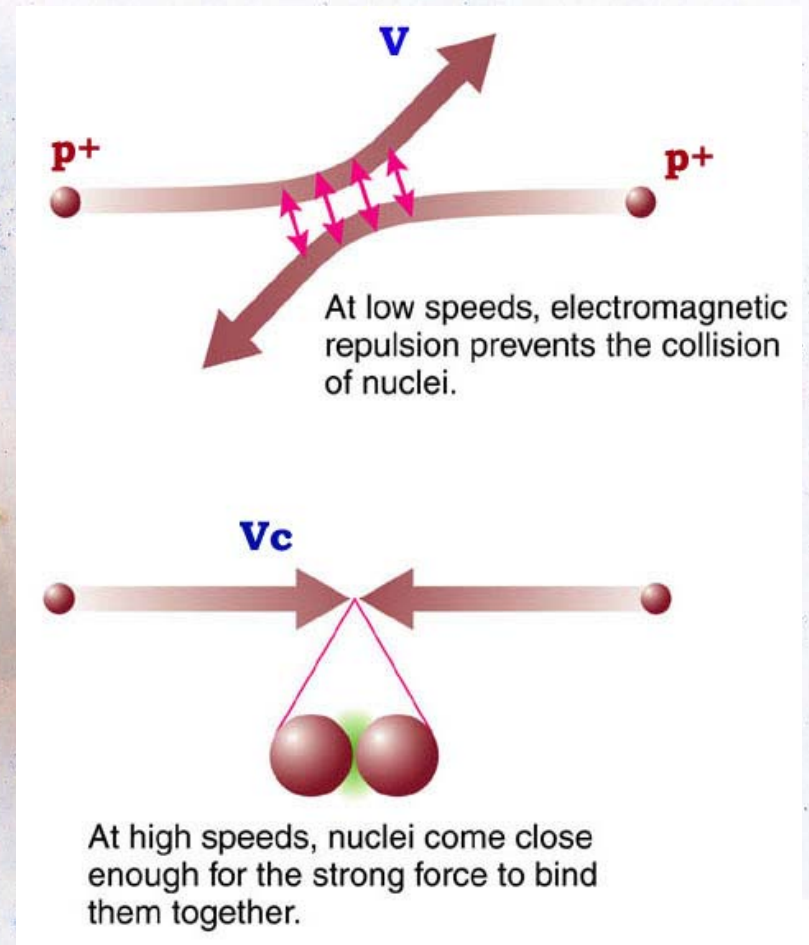


Nuclear Fusion

Protons are positively charged
Particles that are repelled due to
Coulomb force

At high enough speeds particles
Come close enough to FUSE

Nuclear fusion reaction:



High speed of particles = High temperature of gas

The Sun

The Sun is the main sequence star, one of typical stars in our Universe.

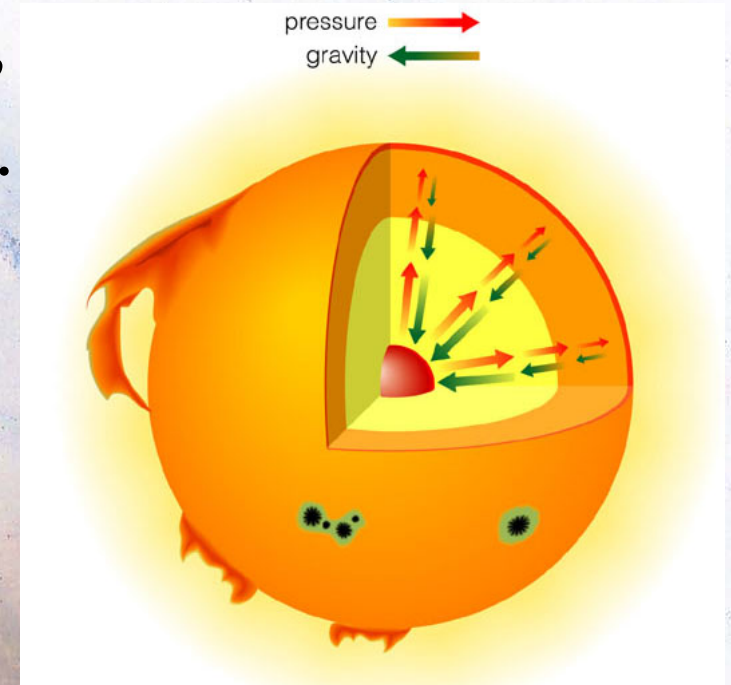
Gravity tends to compress
Compression leads to heating

Core Temperature: 15×10^6 K

Hydrogen ignites into nuclear fusion:

Energy release in the core

Fusion creates pressure to compensate gravity

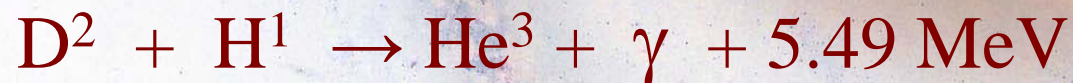
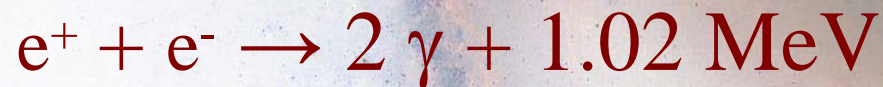


“The theory of stellar nucleosynthesis”

PP chain

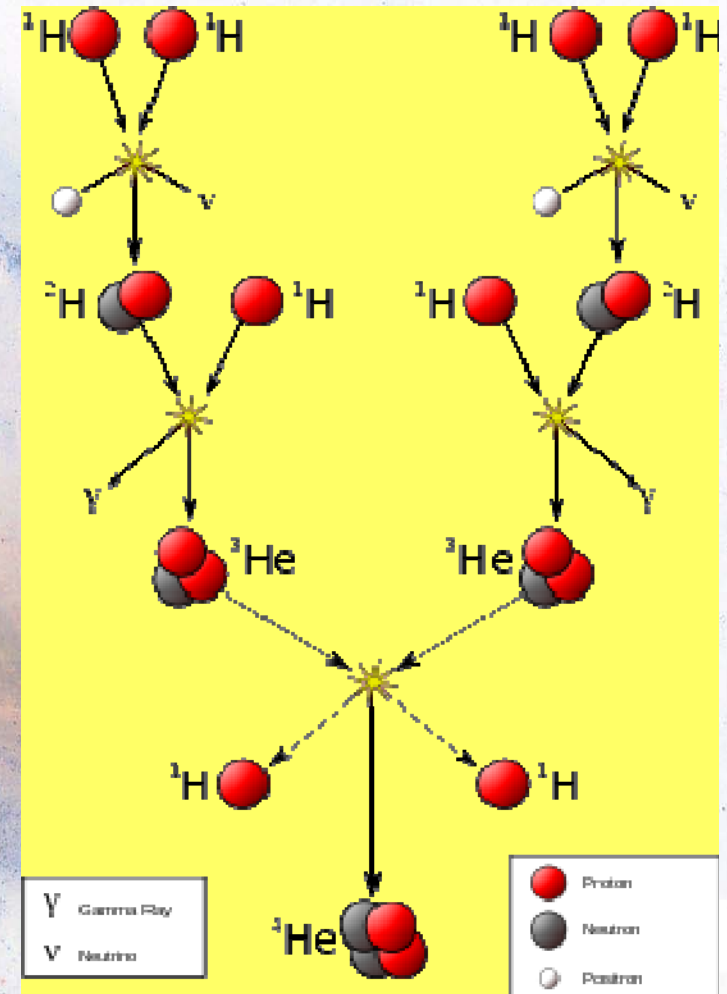
proton-proton chain reactions

$$T = 5 - 15 \times 10^6 \text{ K}$$



$$\Delta E = 0.007 mc^2$$

Energy release: ~ 0.7%



CNO cycle

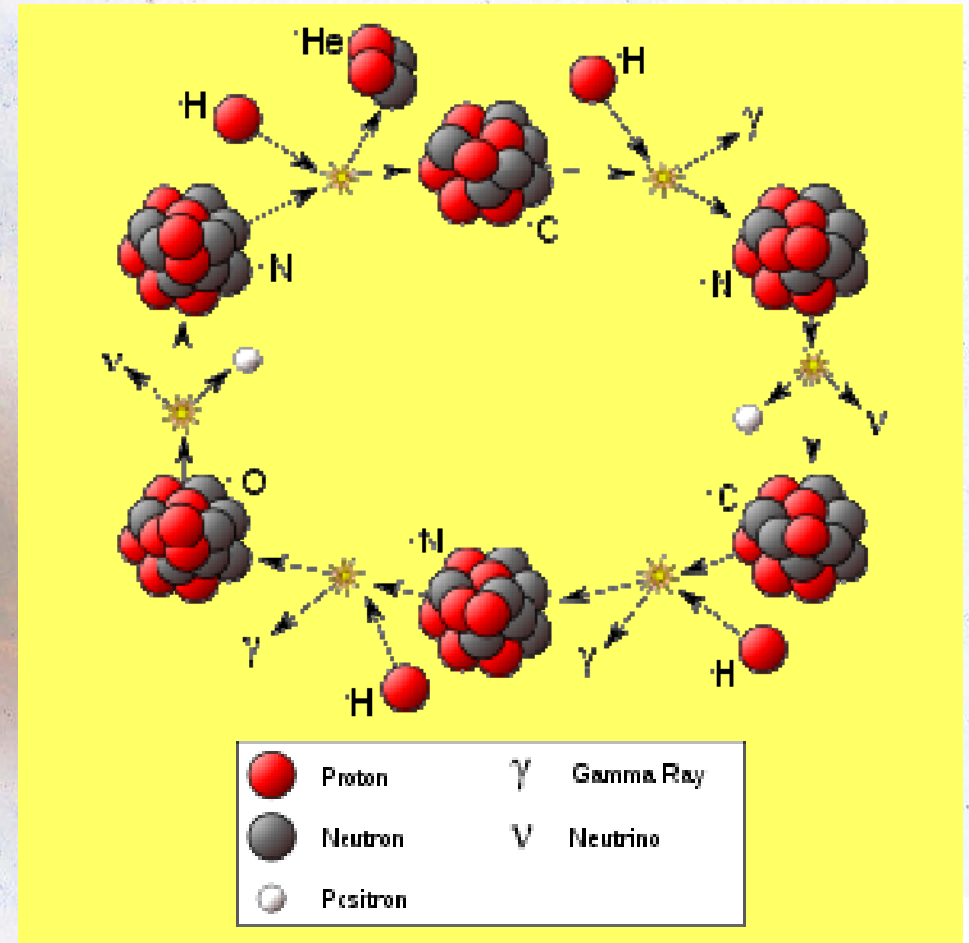
(Bethe-Weizsäcker-cycle)

$M > 1.5 M_{\text{Sun}}$

$T > 20 \cdot 10^6 \text{ K}$

$C + N + O = F + He + \textit{energy}$

Energy release: ~ 0.71%



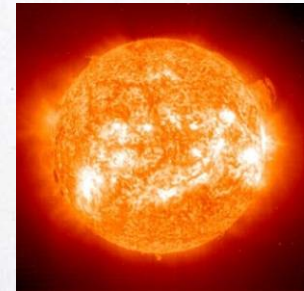
Energy production in massive stars is slightly more effective

Stars in the Sky

Moon, Planets: Reflect Solar Radiation

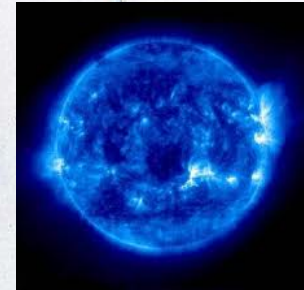


Sun and smaller stars: pp chain



Stars bigger than Sun: CNO cycle

($> 1.5M_{\text{sun}}$)



Galaxies:

Billions of stars often rotating around a common center;

- Huge combined luminosity



Active Galactic Nuclei

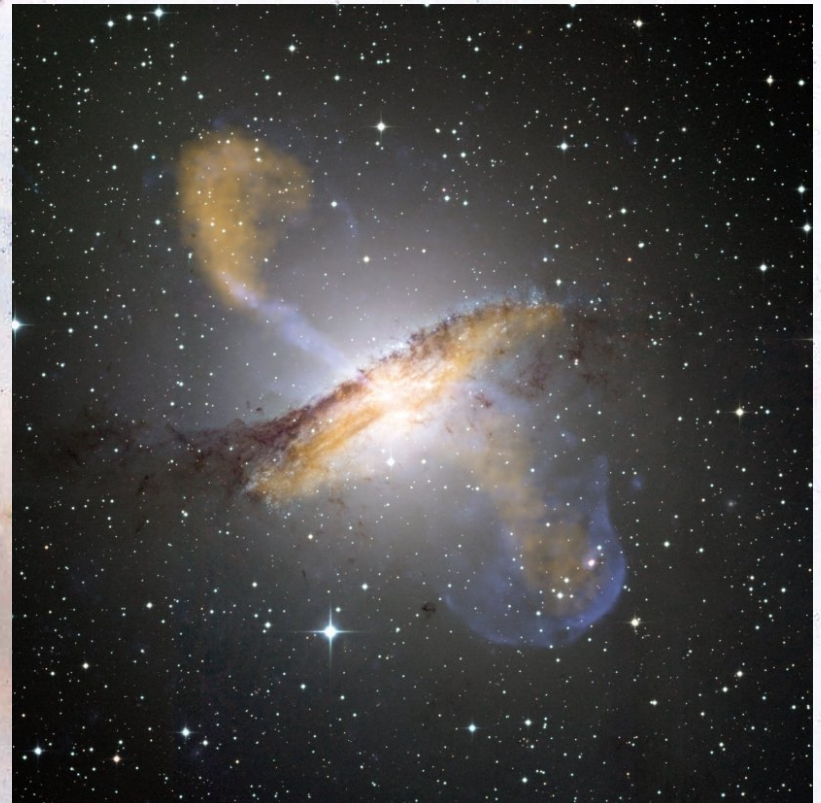
Central parts of some galaxies are extremely luminous

We can estimate the mass of central part by its Gravitational attraction

Non-stellar Luminosity

Can not be explained by simple combination of billions of stars

Energy production rate is much higher than 0.7% limit set by nuclear fusion.



Gravitational Energy

$$E_g = m g h$$

Energy of waterfall:

Gravitational energy



Kinetic Energy



Thermal Energy



Waterfall height:

100m

Water temperature increase:

+0.24 °C

Accretion

Gravitational energy in the Universe

Gravity force: $F_g = GMm/r^2$

Gravity potential: $U = GMm/r$

Energy release: $\Delta E = \Delta U$

How to increase the energy output from mass **m**?

1. Increase **M**
2. Decrease **r**

m mass falling on gravitationally compact object

Accretion

Mass accretion on a black hole

Gas rotating and spiraling down to the center heats up closer to black hole



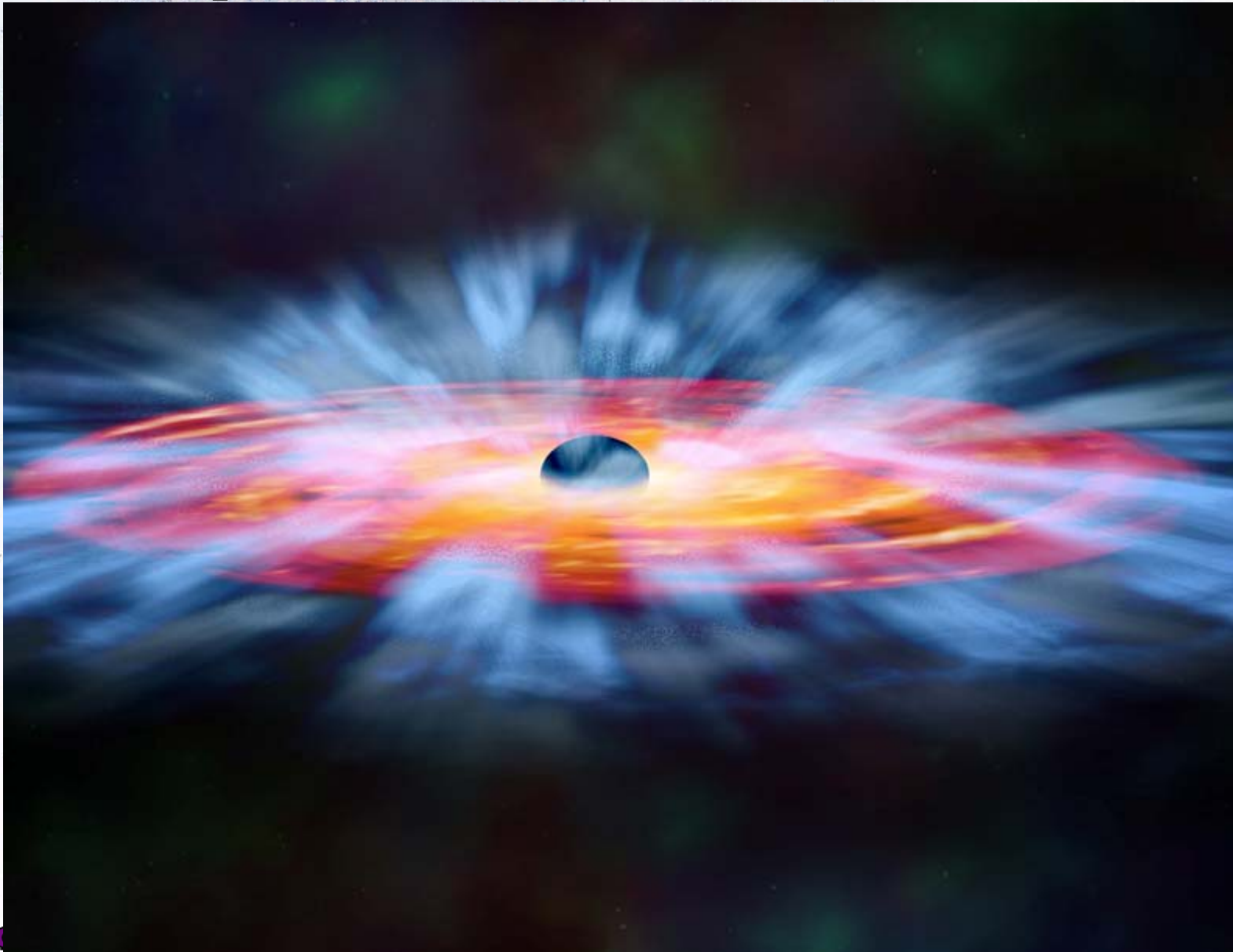
Thermal emission provides huge luminosity (radiation area outside the event horizon)

Energy release: ~ 10 – 20 %

Accretion is the efficient in converting falling mass into the energy

Accretion

Artist's impression of the BH accretion luminosity



AGN Luminosity

Extremely bright galaxy nuclei: $P \sim 10^{37} \text{ W}$

Mechanism: Accretion

Super Massive Black Hole in the center of galaxy

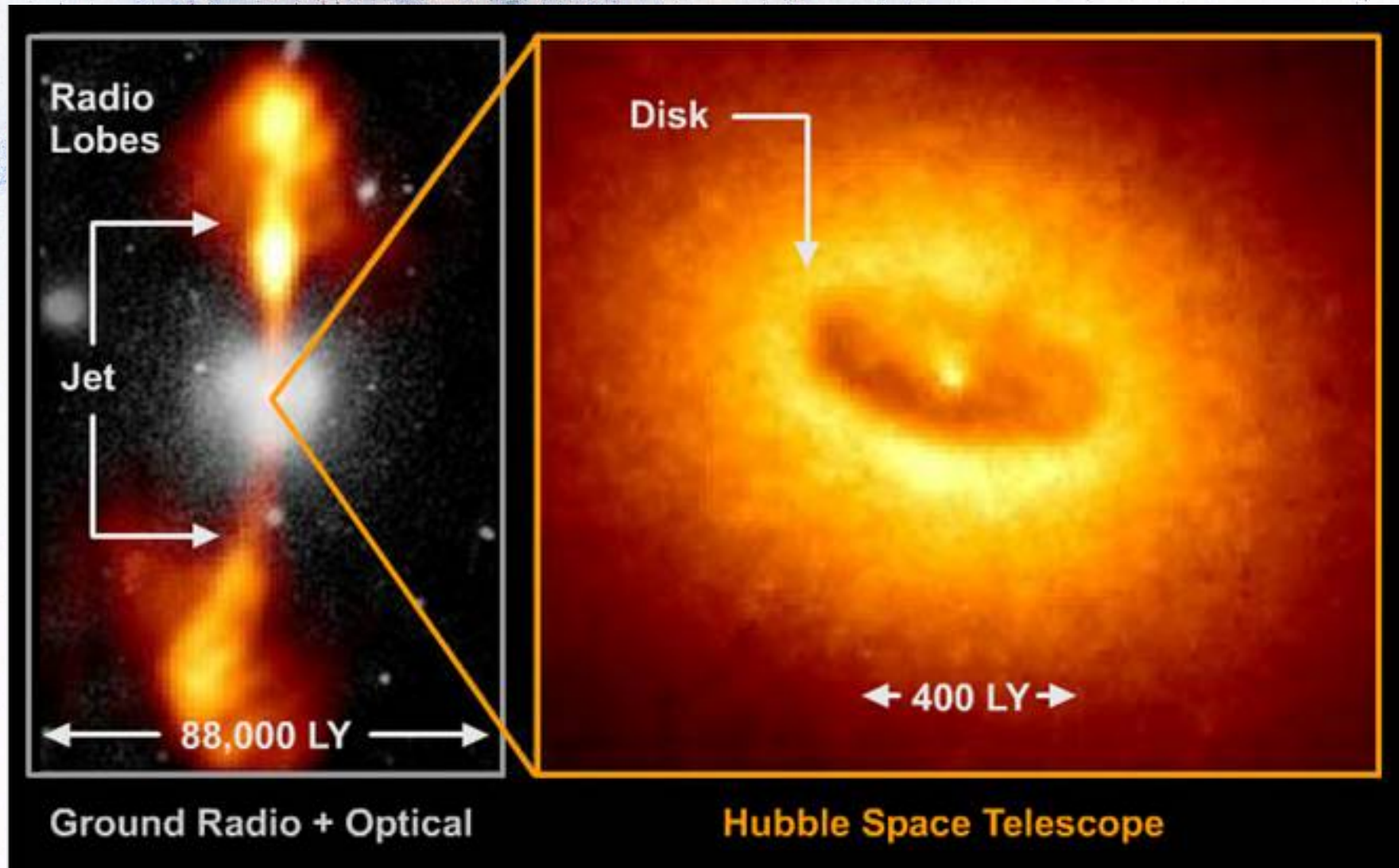
$$M_{\text{BH}} \sim 10^9 M_{\text{Sun}}$$

Mass accretion rate (\dot{m}): $1 M_{\text{Sun}} \text{ in } 1 \text{ year}$

Most Luminous objects in the Universe

Accretion

AGN Observations (Galaxy NGC 4261)



1.) FILL QUIZ !

2.) QUESTIONS? (2:00 pm)

