Lecture 9

Grids

Static Grids:

Uniform Grids;
- Cartesian
- Curvilinear

Non-Uniform Grids;

Irregular Grids;

Dynamic Grids:

AMR (Adaptive Mesh Refinement)

ასტროფიზიკის და პლაზმის ფიზიკის ამოცანების მოდელირება
ალ. თევზაძე (2011)

Grids

Grid Geometry Dependences
- Form of the Equations
- Numerical Differentiation
- Boundary Conditions

Why?

- CPU time
- Memory
- Accuracy

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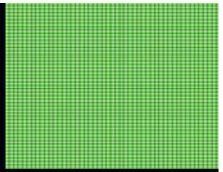
ალ. თევზაძე (2011)

Uniform Grid



Numerical Differentiation

Forward

y'(x) = [y(x) – y(x-h)]/h

Backward

y'(x) = [y(x+h) – y(x)]/h

Three point derivative

$$\frac{f(x+h)-f(x-h)}{2h}$$

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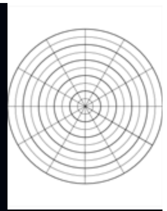
Uniform: Polar



Curvilinear coordinates

1.) Euler Equation: additional terms

$$\frac{\partial V_r}{\partial t} + (\mathbf{V}\cdot\nabla)V_r - \frac{V_\phi^2}{r} = -\frac{1}{\rho}\frac{\partial P}{\partial r} - \frac{\partial \Phi}{\partial r},$$

$$\frac{\partial V_\phi}{\partial t} + (\mathbf{V}\cdot\nabla)V_\phi + \frac{V_r V_\phi}{r} = -\frac{1}{\rho r}\frac{\partial P}{\partial \phi},$$

$$\frac{\partial V_z}{\partial t} + (\mathbf{V}\cdot\nabla)V_z = -\frac{1}{\rho}\frac{\partial P}{\partial z} - \frac{\partial \Phi}{\partial z},$$

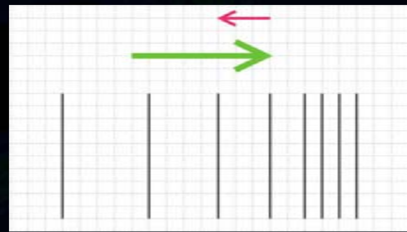
2.) Curvilinear Derivatives

$$(\mathbf{V}\cdot\nabla) \equiv V_r\frac{\partial}{\partial r} + \frac{V_\phi}{r}\frac{\partial}{\partial \phi} + V_z\frac{\partial}{\partial z}.$$

Non-Uniform Grids

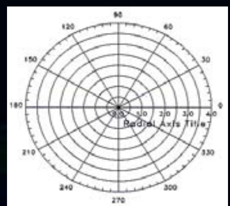
Numerical Derivative: $\frac{f(x+h_2)-f(x-h_1)}{h_1+h_2}$

Stretching factor:
limited

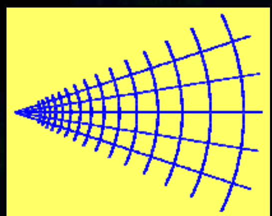


Non-Uniform Curvilinear

ϕ/r assymmetry



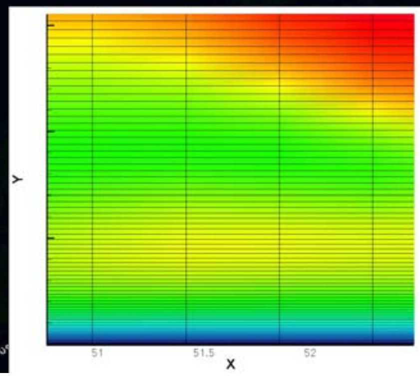
Radial stretching



Chebyshev Grid

Chebyshev polynomials

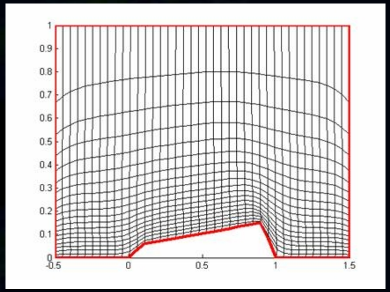
Spectral Method on non-uniform grids



Irregular Grids

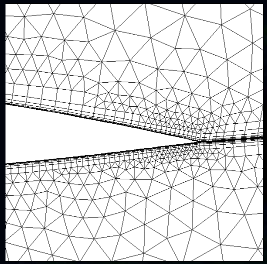
Grid Generation:
Problem-specific grid geometry

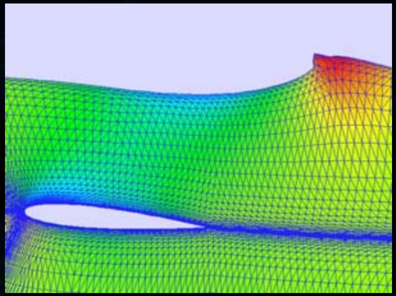
non-regular grids



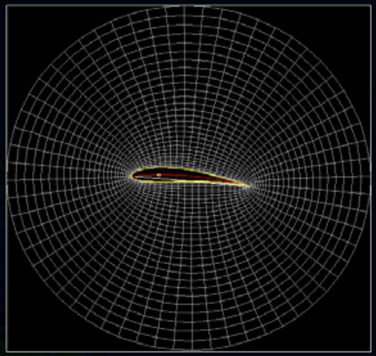
Complex boundaries

varying cell geometry
(industry)





Irregular polar non-uniform



AMR

Adaptive Mesh Refinement:

Multi-scale problems
Scale 1: ***L1 ~ 0.01 m***
Scale 2: ***L2 ~ 100 m***

Δx ~ 0.001m
L ~ 1000m
N ~ 10^6

2D: ***N^2 ~ 10^12***
3D: ***N^3 ~ 10^18***

AMR

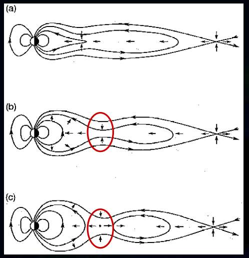
Dynamic mesh geometry:

Adaptation to problem

- Magnetic reconnection
- Self gravity
- Multiscale phenomena

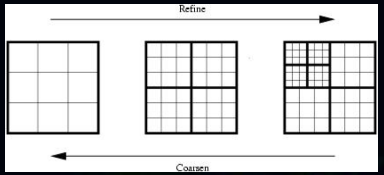
AMR problems

Reconnection

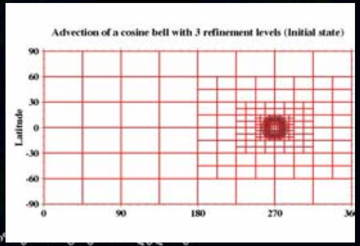


Gravitational clustering

AMR



Refinement levels

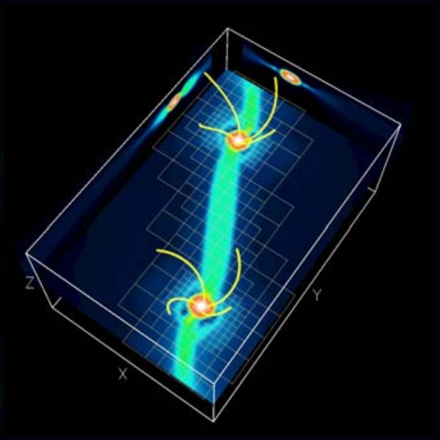


AMR

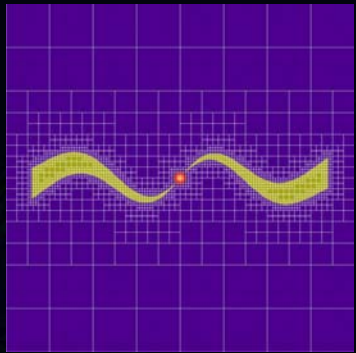
Star formation:

Density structure of a barotropic collapse with magnetic field, computed with NIRVANA3. adaptive mesh refinement and self-gravitation.

U. Ziegler (2005)



AMR in action



Summary

Static Cartesian Grid (simple, fast)
- Polar, Spherical (rotation, axial symmetry)
- Non-Uniform (increasing resolution, static setup)
- Chebishev (Boundary effects, spectral)

AMR (Multiscale problems)

Direct comparison:

More Complex, More CPU, Less Memory

Performance = Balance (CPU,Memory)

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end

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